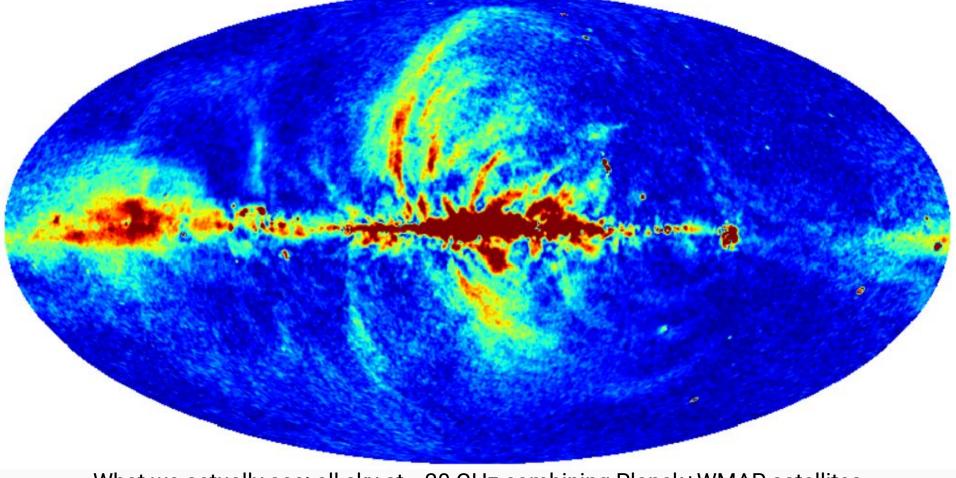


Example of what we want to observe: the cosmic Microwave Background (full sky, in intensity, from the Planck Satellite—polarisation fainter)

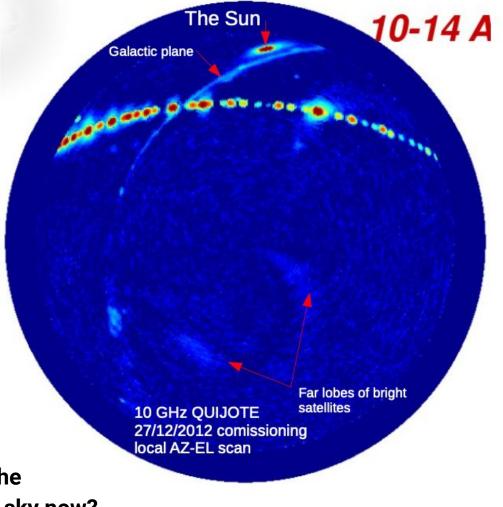


What we actually see: all sky at ~20 GHz combining Planck+WMAP satellites Large scale polarised Synchrotron emission (similar at high freq from dust emission)

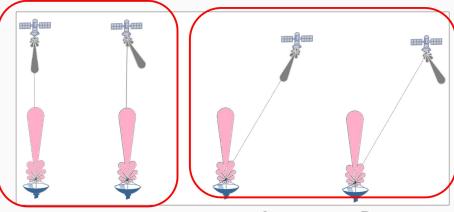
What we really see

- 10−14 GHz local sky from Tenerife
- The Sun, our Galaxy, ...
- Geostationary satellites!
 - Brighter than the sun!
- Satellite signals reflected from the edges of the dish
 - Using special telescopes to minimise sidelobes! Adding extra baffles reduced this.
- This was 2012...
- Satellite numbers now doubled
- Restarting observations this year...

Will we see the equivalent of the geostationary band everywhere in the sky now?



What we might see?



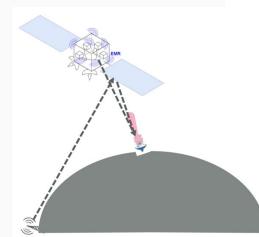
High power Prx

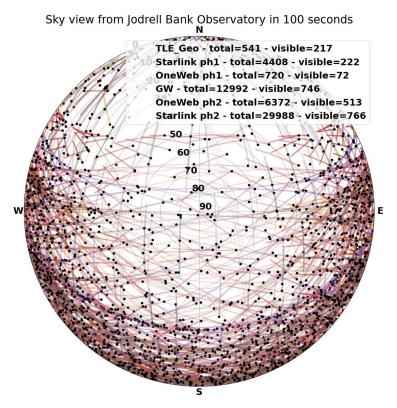
Direct transmissions, in main beam or sidelobes

Unintentional emission from satellite or reflections

Satellites everywhere in the sky all of the time

Low power Prx





Figures courtesy Federico Di Vruno

Why is there a problem (part 1)?

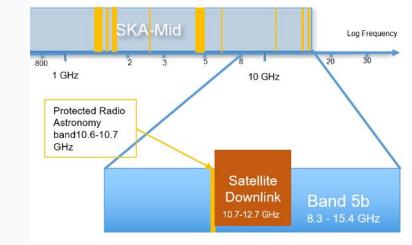
- Radio frequencies regulated by ITU-R
- Includes protected radio astronomy bands
 - Mostly for spectral lines, narrow bands
- Astronomy is 'passive' normally no transmissions to detect faint signals
- Most other uses are 'active' transmissions with high power
- Transmissions can also 'leak' into unintended bands
 - Has happened with constellations like Iridium
 - 10.6-10.7GHz band protected, 10.7+ used for Starlink etc...
- Local radio quiet zones: no transmitters/mobiles, sometimes cars/cameras banned!
 - Basically: have telescopes as far away from people, and in protected areas!

Why is there a problem (part 2)?

- We observe broad frequency ranges (reserved bands very narrow)
 - Sensitivity goes as sqrt(bandwidth x integration time)
 - Can only see some phenomenon at some freqs (spectral lines, spinning dust, ...)
- We survey large sky areas to observe earliest moments (largest scales) in the Universe
- We need high sensitivity to observe very faint signals
 - Using 10,000+ pixels (large focal planes—unprecedented at radio frequencies!)
 - Observe for multiple years
 - Even signals in sidelobes can cause significant problems
- Previously could avoid interference by going to remote parts of the planet
 - Radio quiet zones don't include satellites...
 - >10 GHz frequencies mostly clear free of interference—until now!

What will be impacted?

- Big radio projects running or being planned:
- CMB: CMB-S4 (\$700m), GroundBIRD (>\$3m),
 QUIJOTE (>€10m), ...—Single dishes with
 0.1% sidelobes: still see satellites in sidelobes



- Interferometers: Square Kilometer Array (\$1b?), ALMA (\$1.4b), VLA (>\$100m), eMerlin (> £50m), VLBI (>\$50m) see satellites at 500 km with 1000 km separated telescopes!
 Super-bright signals (even out of band) can make receivers go non-linear, losing observation time (and maybe even damaging receivers)
- Single dishes: Sardinia (>€70m), Green Bank (>\$95m), Effelsberg (>€50m)
- Multi-frequency surveys (separate different astronomical components via frequency)

(All cost estimates are approximate based on public info)

What do signals look like? (small dish)

Presumed **Starlink** satellites

Single pointing direction

11.2-12.2GHz data

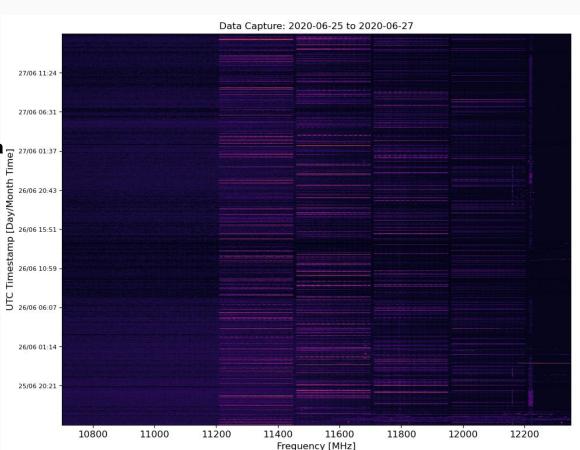
badly

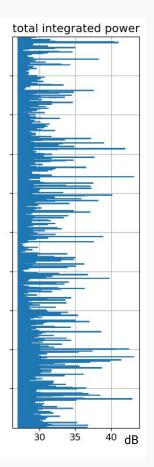
contaminated
(1GHz band!)

Variable (due to satellite movement)

This is over 1.5 days

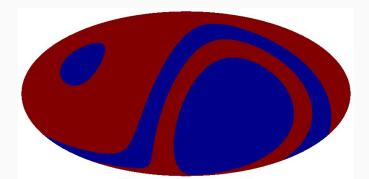
Thanks to Federico Di Vruno (SKAO)

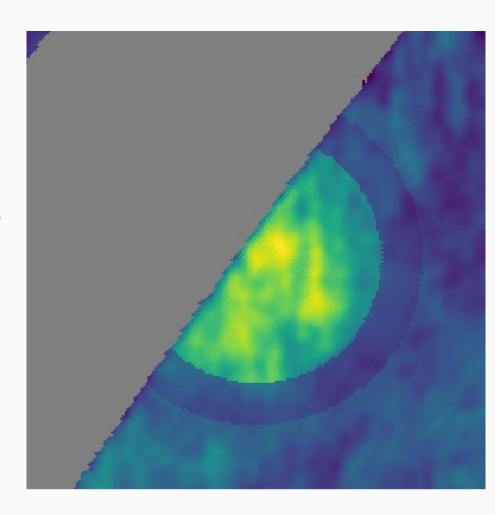




Real on-sky impact?

- Zeta Ophiuchus
- Complex to understand (mix of free-free, sync, maybe AME, acts as Faraday screen, ...)
- Seems to have more high-freq signal than expected (but not spinning dust?)
- Multi-freq component separation would be nice
- But we only see half the source with QUIJOTE—thanks GEO sats! Have to mask ~10° around dec=0° (~100% data loss).
- (Maybe could be filled in with a southern telescope—but \$\$\$!)





What astronomers can do (1/2)

- Use digital backends to split broadband receivers into narrow frequency channels
 - Can be selective about received frequencies within bandwidth
 - Expensive! Particularly at high frequencies.
 - Can only be used for some receivers (not bolometers/KIDs/etc.)
 - Depending on satellite transmissions, may still lose a lot of bandwidth
- Avoid looking at satellites
 - Need to predict where satellites will be, and actively steer around them.
 - Difficult for survey telescopes scanning at fixed elevations
 - Sidelobes still an unavoidable issue

What astronomers can do (2/2)

Early observations

- Feedback to satellite operators to minimise bandpasses/sidelobes
- Knowing out-of-band transmissions from satellites particularly important (e.g., see Iridium transmissions in protected radio astro band!)
- Share observations within radio astro community

Observe for longer

- Estimates depend on bandwidth and time lost, but perhaps 50% longer.
- Construction costs same—but more maintenance/running costs.
- Huge impact on costs of observing and analysis/scientist time (varies between projects, can easily be >50% of telescope costs—or more!)

What we need others to do

- Narrower frequency bands and strict control of out-of-band signals (& tested!)
 - Best thing that can be done—but remember out-of-band leakage!
 (simulated impact from GPS on 21cm: Harper & Dickinson, arXiv:1803.06314)
- Fewer satellites (and less duplication of coverage)
 - Turn them off when passing over all radio telescopes (pro & amateur)
 - Current situation time-consuming + painful, ~100,000 satellites impossible!
- Steerable beams? (Avoid telescopes, or just cause worse sidelobes?)
 - Need to confirm how this impacts telescopes in reality.
- Fainter transmissions (lower power—and stable!)
- More publicly available information (via SatHub?)
 - Bandpasses/transmission frequencies (from measurements)
 - Accurate position predictions

New IAU Centre for the Protection of the Dark and Quiet Sky from Satellite Constellation Interference.

Led by **NOIRLab** (USA) and **SKAO** (UK), with 'Contributing Members' and 'Affiliated Members' Now **open for membership!**.

Director: Piero Benvenuti. Co-directors by Connie Walker & Federico Di Vruno. Four hubs:

- SatHub (leads: Meredith Rawls, Mike Peel)
 - Collection & analysis of satellite observations
 - Software tools
 - Training + outreach
- Policy (leads: Andrew Williams, Richard Green)
 - Coordinate policy action & diplomacy
- Community Engagement (leads: John Barentine, Jessica Heim): beyond professional astronomers
- Industry and Technology (leads: Chris Hofer, Tim Stevenson): engaging industry









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