1.2 Some Quantum Mechanics

1.2.1 The Interpretation of the Wave Function

Before, used $\rho(\underline{x},t)d^3x = |\psi(\underline{x},t)|^2 d^3x$ to get the probability of finding a particle in volume d^3x at \underline{x} at time t. (Assuming normalized ψ .) This requires:

- 1. $\rho(\underline{x},t) = |\psi|^2 \ge 0$
- 2. For a single particle, $\int \rho(\underline{x},t)d^3x = \int |\psi(\underline{x},t)|^2 d^3x = 1$ at any time *t*, and hence $\frac{\partial}{\partial t} \int \rho(\underline{x},t)d^3x = \frac{\partial}{\partial t} \int |\psi(\underline{x},t)|^2 d^3x = 0$ (1) to conserve probability.

Hence we need to show:

$$\frac{\partial \rho}{\partial t} + \underline{\nabla} \cdot \underline{j} = 0$$

(Equation of continuity), where j is a corresponding current density.

Consider the Schrödinger Equation.

$$i\hbar\frac{\partial\psi}{\partial t} = -\frac{\hbar^2}{2m}\nabla^2\psi + V(\underline{x})\psi \quad (1)$$

Taking its conjugate:

$$-i\hbar\frac{\partial\psi^{*}}{\partial t} = -\frac{\hbar^{2}}{2m}\nabla^{2}\psi^{*} + V(\underline{x})\psi^{*} (2)$$

Multiply (1) by ψ^* , and then subtract (2) multiplied by ψ .

$$i\hbar\left(\psi*\frac{\partial\psi}{\partial t}+\frac{\partial\psi*}{\partial t}\psi\right)=-\frac{\hbar^2}{2m}\left(\psi*\nabla^2\psi-\psi\nabla^2\psi*\right)$$

i.e.

$$\frac{\partial}{\partial t}(\psi^*\psi) = \frac{i\hbar}{2m}\underline{\nabla}\cdot(\psi^*\underline{\nabla}\psi - \psi\underline{\nabla}\psi^*)$$

which is of the desired form, where $\rho(\underline{x},t) = |\psi(\underline{x},t)|^2 \ge 0$ and

$$\underline{j} = \frac{i\hbar}{2m} (\psi \underline{\nabla} \psi^* - \psi^* \underline{\nabla} \psi).$$

The continuity equation implies that

$$\frac{\partial}{\partial t} \int_{all \ space} \left| \psi(\underline{x}, t) \right|^2 d^3 x = 0$$

So the conditions for the Born interpretation are satisfied.

1.2.2 Minimal EM Interactions

Point particle with mass m and charge q at position \underline{x} in an EM field (ϕ, \underline{A}) . We want to know what the QM equation of motion for this particle is.

Start from the classical Hamiltonian,

$$H\left(\underline{x},\underline{p}\right) = \frac{1}{2m} \left(\underline{p} - \frac{q}{c}\underline{A}\right)^2 + q\phi \quad (1)$$

To verify this, check Hamilton's Equations of Motion (using equations for many particles in places).

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1.
$$\dot{x}_i = \frac{\partial H}{\partial p_i} = \frac{1}{2m} 2 \left(p_i - \frac{q}{c} A_i \right)$$

i.e. the conjugate momentum $p_i = m\dot{x}_i + \frac{q}{c}A_i$, i.e. $\underline{p} = m\underline{\dot{x}} + \frac{q}{c}\underline{A}$ (2)

2. $\dot{p}_i = -\frac{\partial H}{\partial x_i}$ (somewhat tricky derivation) $m\underline{\ddot{x}} = q\underline{E} + \frac{q}{c}\underline{v} \times \underline{B}$

Note that (1) + (2) implies that the energy is

$$H\left(\underline{x},\underline{p}\right) = E\left(\underline{x},\underline{\dot{x}}\right) = \frac{1}{2}m\underline{\dot{x}}^{2} + q\phi$$

Note that there is no contribution from the magnetic field, as it is at right angles to the particle and hence can't do any work on it.

Schrödinger Equation:

$$H\left(\underline{x},\underline{\hat{p}}\right)\psi = i\hbar\frac{\partial\psi}{\partial t}$$

where $\underline{p} \rightarrow \underline{\hat{p}} = -i\hbar \nabla$ (position operator is still \underline{x}).

$$-\frac{\hbar^2}{2m}\left(\underline{\nabla}-\frac{iq}{\hbar c}\underline{A}\right)^2\psi(\underline{x},t)=i\hbar\left(\frac{\partial}{\partial t}+\frac{iq}{\hbar}\phi\right)\psi(\underline{x},t)$$

which is related to the "free" equation

$$\frac{\hbar^2}{2m}\nabla^2\psi = i\hbar\frac{\partial\psi}{\partial t}$$

by the minimal substitution

$$\frac{\nabla}{\partial t} \rightarrow \frac{\nabla}{\partial t} - \frac{iq}{\hbar c} \frac{A}{\Delta t}$$
$$\frac{\partial}{\partial t} \rightarrow \frac{\partial}{\partial t} + \frac{iq}{\hbar} \phi$$

Or in 4D notation,

$$\partial_{\mu} \rightarrow \partial_{\mu} + \frac{iq}{\hbar c} A_{\mu}(x)$$
 (3)

(This combines the two substitutions before.)

The same argument works in the relativistic case.

$$H = \sqrt{\left(c\underline{p} - q\underline{A}\right)^2 + m^2 c^4} + q\phi$$

Hence the "Schrödinger equation" is

$$\left(\sqrt{-\hbar^2 c^2 \left(\underline{\nabla} - \frac{iq}{\hbar c}\underline{A}\right)^2 + m^2 c^4}\right) \psi(\underline{x}, t) = i\hbar \left(\frac{\partial}{\partial t} + \frac{iq}{\hbar}\phi\right) \psi(\underline{x}, t)$$

Again this is obtained from the free particle case by (3). Note that in the non-EM case, the first bit would form $\sqrt{p^2c^2 + m^2c^4}$.

Problem:- how do we interpret $\sqrt{}$, or in the free case $\sqrt{-\hbar^2 c^2 \nabla^2 + m^2 c^4}$? We will come to this later.

1.3 Natural Units

Use natural units, such that $\hbar = c = 1$. So the unit of speed is the speed of light, and the unit of momentum is \hbar .

It's easy to put this in; the tricky bit is to get back to ordinary units. So use dimensions to restore \hbar , c, and then use

$$\hbar = 6.582 \times 10^{-22} MeV \text{ sec}$$

 $\hbar c = 1.973 \times 10^{-13} MeV m$.

(See Martin & Shaw PP, section 1.5)

2. The Klein-Gorden Equation

Using free space, so $(\phi, \underline{A}) = 0$, and natural units. So the Schrödinger Equation is

$$H\phi(\underline{x},t) = i \frac{\partial\phi(\underline{x},t)}{\partial t} (1)$$

where ϕ is now the wave function, and $H = \sqrt{-\nabla^2 + m^2}$ (2). What does this mean? In the Klein-Gordon (KG) equation, we avoid the problem by noting that *H* is independent of *t*, so we can multiply the equation by *H* again.

$$H^{2}\phi(\underline{x},t) = i\frac{\partial}{\partial t}H\phi = -\frac{\partial^{2}\phi}{\partial t^{2}}$$

then using $H^2 = -\nabla^2 + m^2$,

$$\left(\frac{\partial^2}{\partial t^2} - \underline{\nabla}^2 + m^2\right) \phi(\underline{x}, t) = 0 \quad (3a)$$
$$\left(\Box + m^2\right) \phi(\underline{x}, t) = 0 \quad (3b)$$

where $\Box = \partial_{\mu}\partial^{\mu}$. This is called the Klein-Gordon Equation. This is the correct equation, but what is its interpretation?

There are a whole series of problems.

2.1 Negative Energies

KG equation has solutions

$$\phi(\underline{x},t) = e^{i(\underline{p}\cdot\underline{x}-Et)}$$

with $E^2 = \underline{p}^2 + m^2$, i.e. with energies $E = +\sqrt{\underline{p}^2 + m^2} \ge m^2 > 0$, and $E = -\sqrt{p^2 + m^2} \le m^2 < 0$.

So we have a spectrum of energies available at $E > mc^2$ and $E < -mc^2$, with an energy gap in the middle. Classical particles can't jump this energy gap. Quantum mechanics say that interaction can cause quantum jumps releasing or absorbing quanta $\hbar\omega \ge 2mc^2$, either as radiation or as other particles.

So a particle can fall to negative energy states etc. releasing an infinite amount of energy. What stops it?