

3. The Thermal Physics of the Ideal Classical Gas

(Mandl 7.1-7.6, B&S 5.9, 6.5, 7.4, K&K p72-77)

To show amongst other things that

$$pV = nk_B T$$

in the classical limit.

3.1 What is the classical limit?

Typical energy of particles in a classical gas $\sim k_B T = \left\langle \frac{p^2}{2m} \right\rangle$

$$p \sim (2Mk_B T)^{1/2}$$

de Broglie wavelength of a typical particle:

$$\lambda_T = \frac{h}{p} \sim \frac{h}{(2Mk_B T)^{1/2}}$$

It is conventional to insert π .

$$\text{Define } \lambda_T = \text{thermal wavelength} = \frac{h}{(2\pi Mk_B T)^{1/2}}$$

Classical behaviour is seen if the particle spacing $d \sim \left(\frac{V}{N} \right)^{1/3} \gg \lambda_T$ i.e. quantum

effects are unimportant if the particle spacing is much greater than the de Broglie wavelength of a typical particle.

Define $n_Q = \frac{1}{\lambda_T^3}$ to be the quantum concentration.

The classical limit is when $n \ll n_Q$, i.e. at low densities.

So the classical limit is a low density, high temperature approximation.

NB: $\lambda_T \rightarrow \infty$ as $T \rightarrow 0$, so classical behaviour always fails at low temperature.

3.2 Single Particle Partition Function ζ

Particle function $Z \rightarrow$ free energy $F = -k_B T \ln Z \rightarrow$ pressure, entropy, etc.

For one-particle in a box:

$$\zeta = \sum_{\text{all } k \text{ states}} e^{-\frac{\epsilon(k)}{k_B T}}$$

We can use the approximation

$$\sum e^{-\frac{\epsilon(k)}{k_B T}} \rightarrow (2s+1) \int_0^\infty \frac{V k^2}{2\pi^2} dk e^{-\frac{\hbar^2 k^2}{2Mk_B T}}$$

where:

$(2s+1)$ is if the particles have spin. $s=0$ for simplicity.

$$\frac{V k^2}{2\pi^2} dk = \rho(k) dk$$

$e^{-\frac{\hbar^2 k^2}{2Mk_B T}}$ is the Boltzmann distribution.

$$\zeta = V n_Q = \frac{V}{\lambda_T^3}$$

Integral is straightforward – very important.

$$I = \int_0^\infty e^{-\alpha x^2} x^2 dx$$

where α is the constants, and $x = k$.

$$I = -\frac{d}{d\alpha} \left(\int_0^\infty e^{-\alpha x^2} dx \right)$$

Set $y = \sqrt{\alpha}x$, so $dy = \sqrt{\alpha}dx$.

$$I = -\frac{d}{d\alpha} \left(\frac{1}{\sqrt{\alpha}} \int_0^\infty e^{-y^2} dy \right)$$

$$\int_0^\infty e^{-y^2} dy = \frac{\sqrt{\pi}}{\alpha^{3/2}} \text{ (a well-known integral)}$$

$$I = -\frac{d}{d\alpha} \left[\frac{\sqrt{\pi}}{2\sqrt{\alpha}} \right] = \frac{1}{4} \frac{\sqrt{\pi}}{\alpha^{3/2}}$$

In our integral, $\alpha = \frac{\hbar^2}{2Mk_B T}$ and $x = k$. So:

$$\zeta = \frac{V}{2\pi^2} \frac{\sqrt{\pi}}{4} \left[\frac{2Mk_B T}{\hbar^2} \right]^{3/2} = V \left(\frac{Mk_B T}{2\pi\hbar^2} \right)^{3/2} = \frac{V}{\lambda_T^3} = V n_Q$$

$$\int_0^\infty e^{-y^2} dy = \frac{1}{2} \int_{-\infty}^\infty dy e^{-y^2}$$

$$J = \int_0^\infty dy e^{-y^2}$$

$$J^2 = \int_{-\infty}^\infty dy e^{-y^2} \int_{-\infty}^\infty dx e^{-x^2}$$

(x is the same as y ...)

$$J^2 = \int_{-\infty}^\infty \int_{-\infty}^\infty dx dy e^{-(x^2+y^2)} = \int_0^\infty \int_0^{2\pi} r dr d\theta e^{-r^2}$$

$$= 2\pi \int_0^\infty dr r e^{-r^2} = 2\pi \left[-\frac{1}{2} e^{-r^2} \right]_0^\infty$$

$$= \pi$$

So $J = \sqrt{\pi}$.

3.3 – N Particle Partition Function Z.

(M 7.1, B&S 6.5)

$$Z = \sum e^{-E/k_B T}$$

where the sum is over all the allowed microstates of the N particle system.

For distinguishable particles, $Z = 3^N$.

For indistinguishable particles, $Z \neq 3^N$.

Consider two spinless bosons each of which can be in two energy levels 0 and ϵ .

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For one particle,

$$\zeta = \sum e^{-E/k_B T} = e^{-0} + e^{-\epsilon/k_B T} = 1 + e^{-\epsilon/k_B T}$$

Distinguishable:

Particles labeled a and b.

2 particle system has 4 microstates.

$(a, b) = (0, 0), (0, \epsilon), (\epsilon, 0)$ and (ϵ, ϵ) .

These have energies of 0, ϵ , ϵ and 2ϵ respectively.

The partition function is:

$$Z = e^{-0} + 2e^{-\beta\epsilon} + e^{-2\beta\epsilon} = (1 + e^{-\beta\epsilon})^2 = \zeta^2$$

Indistinguishable:

2 particle system has three microstates $(0, 0)$, $(0, \epsilon)$ and (ϵ, ϵ) . The case of $(\epsilon, 0)$ is dismissed as it is indistinguishable from $(0, \epsilon)$.

Energies are 0, ϵ and 2ϵ .

The partition function is:

$$Z = e^{-0} + e^{-\beta\epsilon} + e^{-2\beta\epsilon} = 1 + e^{-\beta\epsilon} + e^{-2\beta\epsilon} \neq \zeta^2.$$

To prove:

For indistinguishable particles in the classical limit $d \gg \lambda_T$ (or $n \ll n_Q = 1/\lambda_T^3$).

$$Z \approx \frac{\zeta^N}{N!} \text{ where } \zeta \text{ the partition function is } = V n_Q = \frac{V}{\lambda_T^3}.$$

A typical particle in a classical gas has $\epsilon \sim k_B T$. The probability of a state being

occupied is equal to $\frac{e^{-\beta\epsilon}}{\zeta} \sim \frac{e^{-1}}{\zeta} \sim \frac{1}{\zeta}$ for $\epsilon \sim k_B T$.

For N particles in a box, the number of particles in a state $\sim \frac{N}{\zeta} = \frac{N}{V n_Q} = \frac{n}{n_Q}$. But in

the classical regime, $n \ll n_Q$, so $\frac{n}{n_Q} \ll 1$. Hence in the classical limit, there is a very

low probability (negligible) of having more than one particle in any k state.

Accessible k states are more common than occupied states.

Each microstate for indistinguishable particles corresponds to $N!$ microstates for distinguishable particles. $N!$ is the number of arrangements of distinguishable particles. Therefore the sum of the allowed microstates $\sum e^{-E/k_B T}$ is reduced by a factor of $N!$ from that of distinguishable particles.

Therefore:

$$Z = \frac{(V n_Q)^N}{N!} \text{ in the classical limit.}$$

3.4 Properties of the Classical Gas

a) *Helmholtz Free Energy* F

$$F = -k_B T \ln Z$$

b) *Internal energy*

$$E = + \frac{\partial}{\partial \beta} \left(\frac{F}{k_B T} \right) = \frac{\partial}{\partial \beta} (-\ln Z)$$

$$\ln Z = N \ln V + N \ln n_Q - \ln N!$$

$$n_Q \sim \frac{1}{\beta e^{3/2}}, \quad \beta = \frac{1}{k_B T}$$

$$\ln n_Q = -\frac{3}{2} \ln \beta$$

$$E = \frac{3}{2} N \frac{\partial \ln \beta}{\partial \beta} = \frac{3}{2} \frac{N}{\beta} = \frac{3}{2} N k_B T$$

as expected from the equipartition function.

c) *Pressure*

$$P = - \left(\frac{\partial F}{\partial V} \right)_{N,T} = k_B T N \frac{\partial \ln V}{\partial V} = k_B T \frac{N}{V}$$

therefore $pV = N k_B T$ the ideal gas law (in classical limits).

For one mole, $N = N_A$, $pV = RT$. $R = N_A k_B$.

d) *Entropy*

$$S = - \left(\frac{\partial F}{\partial T} \right)_{V,N}$$

Use Sterling's Formula $\ln N! = N \ln N - N + \dots$

$$S = N k_B \left[\ln \left(\frac{V}{N} \right) + \frac{3}{2} \ln(T) + \frac{3}{2} \ln \left(\frac{M k_B}{2 \pi \hbar^2} \right) + \frac{5}{2} \right]$$

the Sackur-Tetrode equation.

NB: it fails at low T. Predicts $S \rightarrow -\infty$ as $T \rightarrow 0$. This is due to the failure of the classical approximation as $T \rightarrow 0$.

If we had included spin, there would have been an additional term in the entropy $N k_B \ln(2s+1)$ where $2s+1$ is the spin degeneracy.

Disorder associated with the spins: 2^N arrangements of the N spins for $S = \frac{1}{2}$.

This is the only quantity affected by the spin degeneracy.

e) *Chemical Potential*

$$\mu = \left(\frac{\partial F}{\partial N} \right)_{V,T}$$

$$\mu = -\frac{3}{2} k_B T \ln \left(\frac{N}{V} \left(\frac{2\pi M k_B T}{h^2} \right)^{3/2} \right) = -\frac{3}{2} k_B T \ln \frac{d^2}{\lambda_T^2}$$

This is negative in the classical regime where $d \gg \lambda_T$.

3.5 How good is the classical approximation at STP?

(STP = Standard Temperature and Pressure)

$$p = \frac{N}{V} k_B T, \text{ so } \frac{N}{V} \approx 3 \times 10^{25} \text{ m}^{-3} \text{ at STP.}$$

The typical distance between particles is $d \sim \left(\frac{V}{N} \right)^{1/3} = 3.2 \text{ nm}$

$$\lambda_T = \frac{h}{(2\pi M k_B T)^{1/2}} \sim 0.05 \text{ nm}$$

So $d > \lambda_T$, so the approximation is good. However it is still the same order of magnitude, so there are quantum effects present – very sensitive experiments can detect them. These effects become much larger at lower temperatures.