

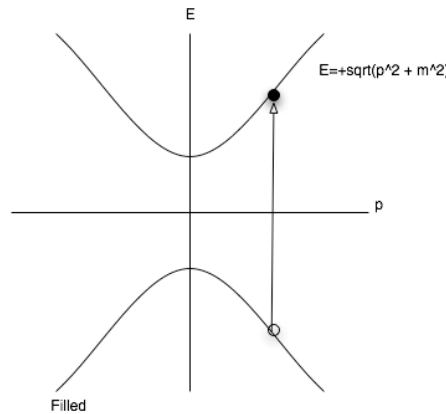
Solutions $E > m$, \underline{p} , $s(= +\frac{1}{2}, -\frac{1}{2})$

Plus solution $-E < -m$, $-\underline{p}$, $-s$.

3.4 Dirac hole theory

Spin-1/2 particles are fermions, and they obey the exclusion principle.

Imagine that the vacuum is a completely filled ‘sea’ of negative energy states.



Positive energy e^- can't fall down to negative energy states because of the exclusion principle.

But you can excite a particle from the sea. The hole looks like an antiparticle, i.e. absence of e^- with $-E, -\underline{p}, -s$ looks like e^+ with E, \underline{p}, s . This is a way to live with negative energy states – for a while!

3.5 Relativistic Covariance

3.5.1 Covariant Notation

Consider the Dirac equation,

$$i \frac{\partial \psi}{\partial t} = [-i \underline{\alpha} \cdot \underline{\nabla} + \beta m] \psi \quad (1)$$

This equation is not obviously covariant.

Multiply by β from the left, and define $\gamma^0 = \beta$, $\gamma^i = \beta \alpha_i$ ($i = 1, 2, 3$).

$$i \gamma^0 \frac{\partial \psi}{\partial x^0} = \left[-i \gamma^i \frac{\partial}{\partial x^i} + m \right] \psi$$

since $\beta^2 = 1$. Writing $\gamma^\mu = (\gamma^0, \gamma^i)$, then (1) becomes

$$(i \gamma^\mu \partial_\mu - m) \psi(x) = 0 \quad (2)$$

where the γ -matrices satisfy

$$\gamma^\mu \gamma^\nu + \gamma^\nu \gamma^\mu = 2 g^{\mu\nu} \quad (3)$$

$$\gamma^{\mu\dagger} = \gamma^0 \gamma^\mu \gamma^0 \quad (4a)$$

$$\text{or } \gamma^{0\dagger} = \gamma^0, \gamma^{i\dagger} = -\gamma^i \quad (4b)$$

which follow from the properties of β , α_i . e.g. $(\gamma^i)^\dagger = (\beta \alpha_i)^\dagger = \alpha_i \beta = -\beta \alpha_i = -\gamma^i$.

It's also useful to introduce the (Dirac) adjoint

$$\bar{\psi}(x) = \psi^\dagger(x) \gamma^0 \quad (5)$$

The charge current density $j^\mu(x) = (\rho, \mathbf{j})$ can be rewritten (exercise)

$$j^\mu(x) = \bar{\psi}(x) \gamma^\mu \psi(x) \quad (6)$$

with the equation of continuity

$$\partial_\mu j^\mu(x) = 0 \quad (7).$$

This all looks covariant – but is it?

If they are valid physical equations, they must be by the principle of special relativity. But what are the transformation properties of ψ ?

3.5.2 Proof of covariance

Consider a Lorentz transformation from $O \rightarrow O'$ such that $x \rightarrow x' = ax$ ($\Lambda_\mu^\nu \rightarrow a_\mu^\nu$)

i.e. $x'^\mu = a^\mu_\nu x^\nu$ (1) where $a^\mu_\nu a_\mu^\sigma = g_\nu^\sigma = \delta_\nu^\sigma$ (2).

What is the transformation law which takes $\psi(x) \rightarrow \psi'(x')$?

Define $S(a)$ by

$$\psi'(x') = S(a)\psi(x) \quad (3)$$

and correspondingly,

$$\psi(x) = S^{-1}(a)\psi'(x') = S(a^{-1})\psi'(x') \quad (4)$$

Then the Dirac equation is covariant and $j^\mu(x) = \bar{\psi}(x)\gamma^\mu\psi(x)$ is a 4-vector, provided:

$$S^{-1}(a)\gamma^\mu S(a) = a^\mu_\nu \gamma^\nu \quad (5)$$

and

$$S^{-1}(a) = \gamma^0 S(a) \gamma^0 \quad (6).$$

To prove this is so, we need to show that the Dirac equation is invariant in form under Lorentz transformation plus (5) and (6), i.e.

$$(i\gamma^\mu \partial'_\mu - m)\psi'(x') = 0 \quad (7)$$

implies

$$(i\gamma^\mu \partial_\mu - m)\psi(x) = 0$$

(or vice versa) where $\partial'_\mu = \frac{\partial}{\partial x'^\mu} = a_\mu^\sigma \partial_\sigma$. (8)

Substitute (3) and (8) into (7).

$$\rightarrow (i\gamma^\mu a_\mu^\sigma \partial_\sigma - m)S(a)\psi(x) = 0$$

Multiply this from the left by $S^{-1}(a)$.

$$\rightarrow \left(iS^{-1}(a)\gamma^\mu S(a)a_\mu^\sigma \partial_\sigma - \underbrace{(S^{-1}(a)S(a))}_{=I} m \right) \psi(x) = 0$$

But by (5), $S^{-1}(a)\gamma^\mu S(a)a_\mu^\sigma = \gamma^\nu \underbrace{a^\mu_\nu a_\mu^\sigma}_{\delta_\nu^\sigma} = \gamma^\sigma$ by (2).

So (9) becomes

$$(i\gamma^\sigma \partial_\sigma - m)\psi(x) = 0$$

as required.

What about $j^\mu(x) = \bar{\psi}\gamma^\mu\psi(x)$?

To show it's a 4-vector, we need to show

$$\bar{\psi}'(x')\gamma^\mu\psi'(x') = a^\mu_\nu \bar{\psi}(x)\gamma^\nu\psi(x).$$

Consider the transformation law of $\bar{\psi}'(x')$.

$$\psi'(x') = S(a)\psi(x) \rightarrow \psi'^{\dagger}(x') = \psi^{\dagger}(x)S^{\dagger}(a)$$

$$\text{Therefore } \bar{\psi}'(x') = \psi'^{\dagger}(x')\gamma^0 = \psi^{\dagger}(x)\gamma^0\gamma^0 S^{\dagger}(a)\gamma^0 = \bar{\psi}(x)\gamma^0 S^{\dagger}(a)\gamma^0$$

$$\text{So } \bar{\psi}'(x') = \bar{\psi}(x)S^{-1}(a) \text{ by (6) (10)}$$

Therefore,

$$\begin{aligned} \bar{\psi}'(x')\gamma^\mu\psi'(x') &= \bar{\psi}(x)S^{-1}(a)\gamma^\mu S(a)\psi(x) \\ &= a^\mu_\nu \bar{\psi}(x)\gamma^\nu\psi(x) \end{aligned}$$

which is equation (9). QED.

Note that (10) also implies $\bar{\psi}(x)\psi(x)$ is a Lorentz scalar, etc.

3.5.3 Lorentz boost

Consider $O \rightarrow O'$

$(x, y) \rightarrow (x', y')$ with difference vt . Then $x' = \alpha x$ where

$$a = \begin{pmatrix} \cosh \omega & -\sinh \omega & 0 & 0 \\ \sinh \omega & \cosh \omega & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

where $\sinh \omega = \frac{v}{\sqrt{1-v^2}}$ and $\cosh \omega = \frac{1}{\sqrt{1-v^2}}$. Then

$$S(a) = e^{\omega\alpha_1/2} = I \cosh \frac{\omega}{2} + \alpha_1 \sinh \frac{\omega}{2} \text{ where we have used the fact that } \alpha_1^2 = 1;$$

$$\alpha_1^{2n} = 1; \alpha_1^{2n+1} = \alpha_1.$$

This satisfies the conditions (5) and (6). (Exercise – verify this).

Note – for particles at rest, $\psi \propto \begin{pmatrix} 1 \\ 0 \\ 0 \\ 0 \end{pmatrix}$ etc., and we can use a Lorentz boost to get

$\psi(v)$.